## The University of Sydney Math3976 (Advanced)

Mathematical Computing: Numerical Methods 2015

Lecturer: D. J. Ivers

## Assignment 2 — Quadrature and Initial Value Problems for Systems of ODE's

Hand-written Questions 1 and 2 due 5pm Thursday 4<sup>th</sup> June. Programming Questions 3 and 4 due 5pm Thursday 4<sup>th</sup> June.

Read the course messages for general information on assignments, the form your Assignment 2 file must take and how to submit it. This assignment is worth 8 marks.

1. The 2-point Gauss rule for the weight  $e^{-x}$  on the interval  $[0,\infty]$  is of the form

$$\int_0^\infty f(x)e^{-x} dx \approx w_1 f(x_1) + w_2 f(x_2).$$

(a) Using the Gram-Schmidt procedure determine the orthogonal polynomials  $\phi_0(x)$ ,  $\phi_1(x)$  and  $\phi_2(x)$  for the weight function  $w(x) = e^{-x}$  in  $0 \le x \le \infty$ . Find also  $A_0$ ,  $A_1$ ,  $A_2$ ,  $\gamma_0$ ,  $\gamma_1$ . Suggestion: show

$$\int_0^\infty x^n e^{-x} dx = n!, \qquad n = 0, 1, 2, 3, \dots.$$

(The integral on the left is actually the Gamma (or factorial) function  $\Gamma(n+1)$ .  $\phi_n(x)$  is the *n*-degree Laguerre polynomial.)

- (b) Determine the nodes  $x_1$  and  $x_2$  for the 2-point weighted Gauss rule from  $\phi_2(x)$ .
- (c) Using the recurrence relation

$$\phi_{k+1}(x) - (a_k x + b_k)\phi_k(x) + c_k \phi_{k-1}(x) = 0$$

given in the notes show that the weights

$$w_k = -\frac{A_{n+1}\gamma_n}{A_n\phi_{n+1}(x_k)\phi'_n(x_k)}$$

can be expressed in the more useful form

$$w_k = \frac{A_n \gamma_{n-1}}{A_{n-1} \phi_{n-1}(x_k) \phi'_n(x_k)} \,.$$

Hence determine the weights  $w_1$  and  $w_2$  for the 2-point weighted Gauss rule.

2. Consider the initial value problem

$$\frac{dy}{dt} = f(t, y), \qquad y(0) = y_0.$$

(a) By expanding the exact solution y = Y(t) of the initial value problem in a two term Taylor series with remainder about  $t = t_{k+1}$ , show that

$$Y_k = Y_{k+1} - hf(t_{k+1}, Y_{k+1}) + \frac{1}{2}Y''(\xi_k)h^2$$
,

where  $Y_k \equiv Y(t_k)$ , etc., and  $\xi_k$  lies between  $t_k$  and  $t_{k+1}$ .

(b) By subtracting the equation derived in Part (a) from the backward-Euler method,

$$y_{k+1} = y_k + h f(t_{k+1}, y_{k+1}),$$

deduce that the global error  $e_k$  in the backward-Euler method satisfies

$$e_{k+1} = \frac{e_k}{1 - hJ} - \frac{Y''(\xi_k)h^2}{2(1 - hJ)},$$

where  $J = f_y(t_{k+1}, \eta_{k+1})$  and  $\eta_{k+1}$  lies between  $y_{k+1}$  and  $Y_{k+1}$ .

(c) Deduce from Part (b) that the backward-Euler method is stable, when applied to well-conditioned problems, if h > 0 or h < 2/J < 0. Thus the step-size h > 0 in the backward-Euler method is unrestricted by stability considerations and can be chosen for accuracy.

3. Attached is a listing of the file ass2.inc which contains an incomplete main program implementing the classical fourth-order Runge-Kutta for solving the second-order initial value problem

$$My'' + ry' + Mg = 0,$$
  $y(0) = H,$   $y'(0) = 0,$ 

for an object of mass M dropped from a vertical height H in a uniform gravitational field g and subject to an air resistance proportional to the velocity. Assume  $M=10\,\mathrm{kg},\ r=0,\ g=10\,\mathrm{m^2\,s^{-1}}$  and  $H=10\,\mathrm{m}$ . In ass2.inc a question mark? indicates a missing single character, three horizontal dots . . . indicates a missing part or the missing whole of a line and three vertical dots indicates a missing block (two or more lines) of code. Save the file to your directory and complete the program to have the following features:

- (a) The problem parameters (mass (for M), r and g in Question 3; mu and mut in Question 4) are input from a file called parameters.dat using a namelist parameters and echoed to the screen. Save the file parameters.dat to your directory.
- (b) The arrays y, yrk4, k1, k2, k3, k4 are declared and dimensioned in the main program.
- (c) The initial time t0, the final time tn, the number of steps n, the output frequency nprint and the vector  $\mathbf{y}_0$  of initial values are input from a file using a namelist run\_data and echoed to the screen. Use the array y for the vector  $\mathbf{y}_0$  of initial values. The name of the data file is input from the workstation with a prompt.
- (d) The values of t and y are written into a file with appropriate headers. An implied do-loop is not used for y. The name of the output file is input from the workstation with a prompt.
- (e) The step-size h is set by t0, tn and n.
- (f) Values of t and y are printed every nprint steps with format descriptors F14.5 and E14.5, respectively.
- (g) The first-order system of ode's is given by the array-valued internal function f. Local arrays y and f\_result are declared and dimensioned in f.
- (h) The fourth-order Runge-Kutta subroutine is rk4 from Set 5.

4. Consider the following Earth-Moon-spaceship problem,

$$x'' = 2y' + x - \frac{\tilde{\mu}(x+\mu)}{r_1^3} - \frac{\mu(x-\tilde{\mu})}{r_2^3},$$
  
$$y'' = -2x' + y - \frac{\tilde{\mu}y}{r_1^3} - \frac{\mu y}{r_2^3},$$

where  $\tilde{\mu} = 1 - \mu$ ,  $r_1^2 = (x + \mu)^2 + y^2$  and  $r_2^2 = (x - \tilde{\mu})^2 + y^2$ . The coordinates of the spaceship are (x,y) in a reference frame rotating with the Earth and the Moon, which are at  $(-\mu,0)$  and  $(\tilde{\mu},0)$ , respectively. The equations have been scaled so that distance is measured in units of the Earth-Moon distance R, time in units of  $\Omega^{-1}$ , where  $\Omega$  is the angular speed of the Moon about the Earth and velocity in units of  $R\Omega$ . Since the mass of the spaceship is negligible compared to the masses of the Earth and Moon, the Earth-Moon-spaceship problem is a case of the restricted three-body problem.

Modify your program from Question 3 to solve the Earth-Moon-spaceship problem. Use the following initial values of the dependent variables

$$x(0) = 0.994$$
,  $y(0) = 0$ ,  $x'(0) = 0$ ,  $y'(0) = -2.0015851$ ,

for  $\mu = 0.012277471$ . Take t0=0, tn=17.5, n=20000 and nprint=50.

- (a) Show numerically that the orbit is periodic and find the period as accurately as possible.
- (b) Find the minimum radius from the Moon to the orbit. Use a block-if in the t\_loop to determine the latest minimum radius r2\_min and the corresponding time t\_min, position and velocity y\_min. Output the results to the screen with appropriate headers after completing the t\_loop. From your output file (output  $t, x, y, x', y', r_2$ ) determine the time(s), position(s) and velocity(ies) of any other points of the orbit, where the minimum radius occurs.
- (c) Use funplot to plot the orbit.

Read the course messages for details on what to hand in.

```
program rk4sysmain
!
! This program solves the initial value problem for a system
! of first-order differential equations using the classical
! fourth-order Runge-Kutta method.
!_____
! Declare problem size:
integer, parameter :: m=?
! Declare program variables:
real :: mass,...
integer :: n, ...
real :: t0, ...
real, dimension(m) :: y, ...
character(len=20) :: datafile, ??????
namelist /???????/ mass,...
??????? /run_data/ t0,...,y
!...Open and enter parameters from file parameters.dat:
open (unit=7, file='parameters.dat', status='old')
read (7,nml=parameters)
close(7)
!...Echo parameters to terminal:
write (*,nml=parameters)
!...Name, open and enter run_data from data_file:
write (*,*) ' Enter run_data file name:'
read (*,10) ????????
10 format (a)
open (????=7, file=datafile, status='old')
read (7,nml=...)
close(7)
!...Echo run_data to terminal:
write (*,???=run_data)
! Name, open and head output file:
write (*,*) 'Enter output file name:'
read (*,10) output
open (unit=8, file=?????, status='unknown')
write (8,30)
30 format('
                                y'' ') !Q1 (adjust spacing if needed)
             t
                     у
!30 format(' t
                                         !Q2
                     X
! Initialise values for DO-loop (y initialised directly by input):
t = \dots
h = \dots
! Initialise min_r2 for Q2 only:
! r2_min = ... !Q2
t_loop: do j= 1, ?
   ! 4th order Runge-Kutta subroutine:
  call rk4
  y = ????
  t = \dots
```

```
! Print solutions every nprint steps:
  if ( mod(j,nprint) == 0 ) then
      r2 = ... !Q2
      write (8,40) t, ..., r2 !Q2
     write (8,40) t, ...
     format (?????, ??????)
40
  endif
  ! The following code is for Question 2 only:
  !r2 = ...
  !Calculate the minimum of r2:
  !if (r2_min > r2) then
  ! r2_min= ...
  !endif
enddo t_loop
! The following \dots is for Question 2 only:
! Output the minimum of r2 and associated values of t,y:
  !
stop
???????
function f(t,y) result(f_result)
real, intent(in):: t
real, dimension(m), intent(in):: y
real, dimension(m) :: f_result
f_result(1) = ...
  !
end ????????????
!-----
subroutine rk4
   ! Purpose: To implement the classical fourth-order Runge-Kutta
   ! method for the system of M first-order differential equations
   ! y'=f(t,y) over one interval of width h. Thus subroutine rk4
   ! approximates y(t+h) with yrk4, given y=y(t).
k1 = h*f(t,y)
k2 = h*f(t + h/2.,y + k1/2.)
k3 = h*f(t + h/2., y + k2/2.)
k4 = h*f(t+h,y+k3)
yrk4 = y + (k1 + 2.*k2 + 2.*k3 + k4)/6.
return
end subroutine rk4
I-----
end ...
```